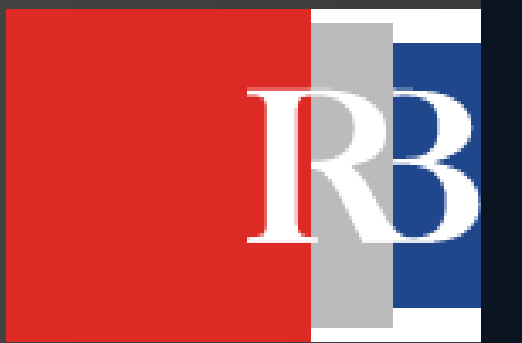


# Neutrino oscillation results of the OPERA experiment in the CNGS beam

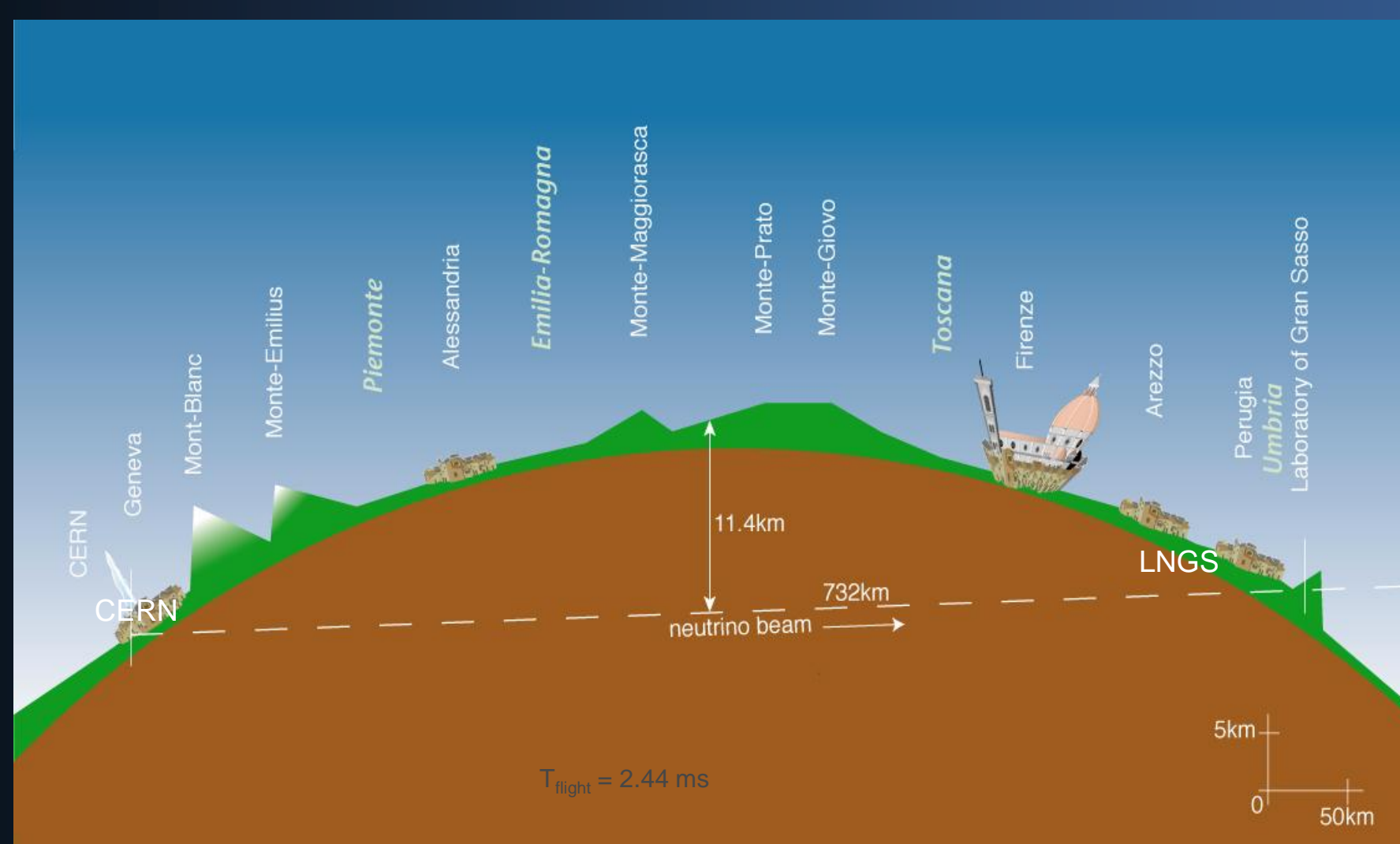
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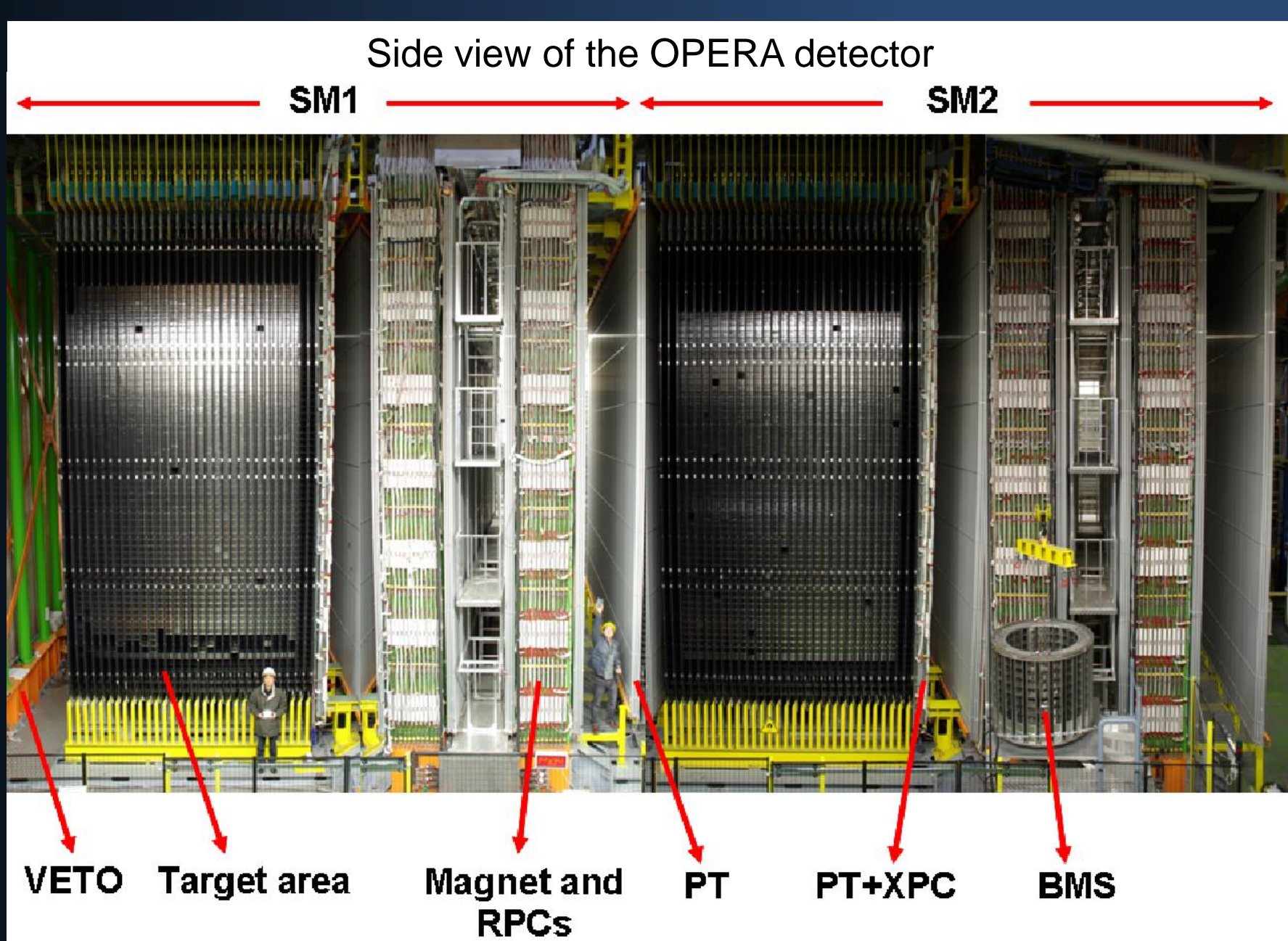


## The OPERA experiment

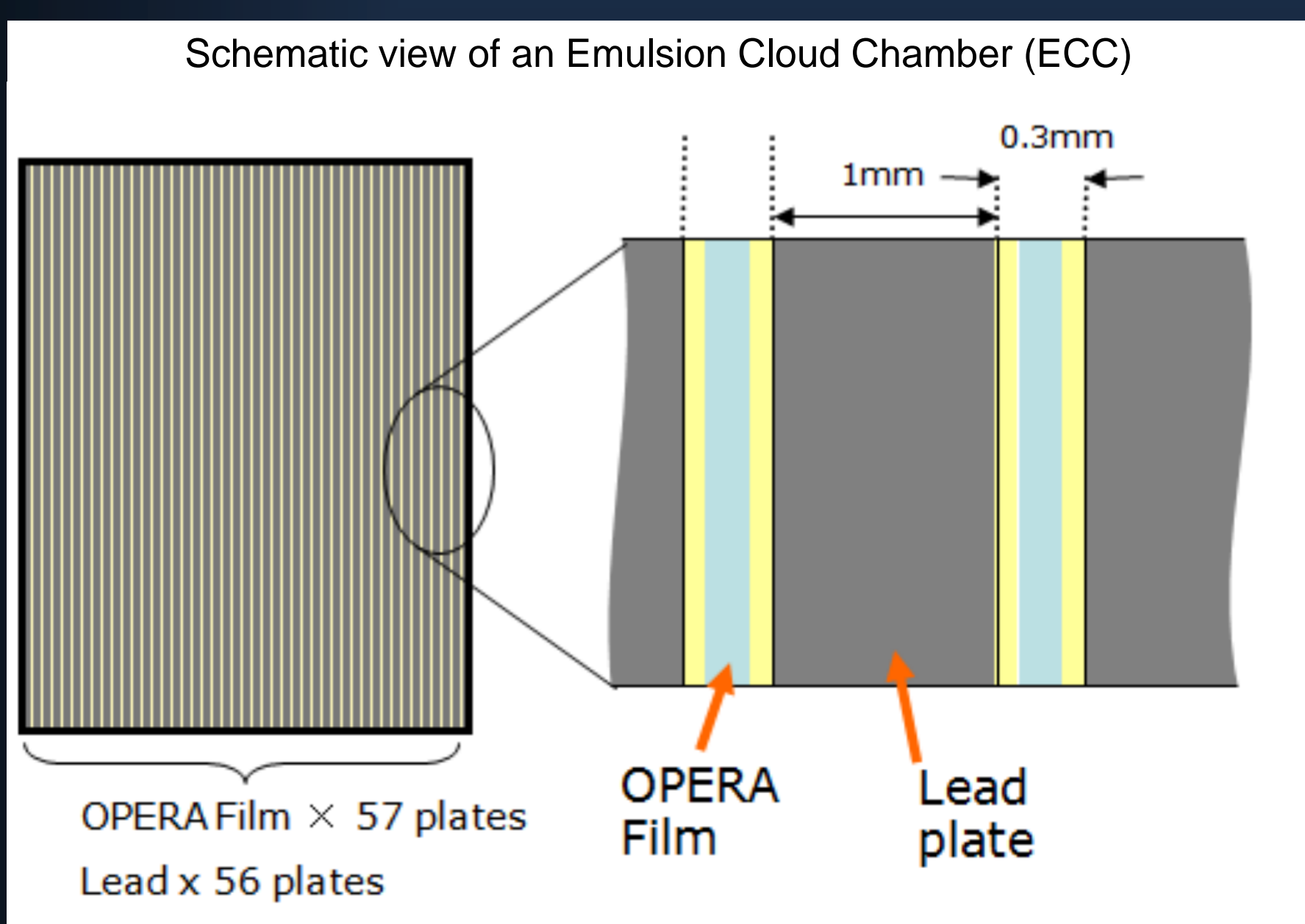
The OPERA experiment [1] was designed to detect tau neutrino appearance in the predominantly muon neutrino CNGS beam [2].



The OPERA detector target consisted of about 150 thousand stacks of lead plates and emulsion films, called Emulsion Cloud Chamber (ECC) elements, embedded between scintillator planes. The detector also featured two muon spectrometers, used for momentum reconstruction of charged particles.

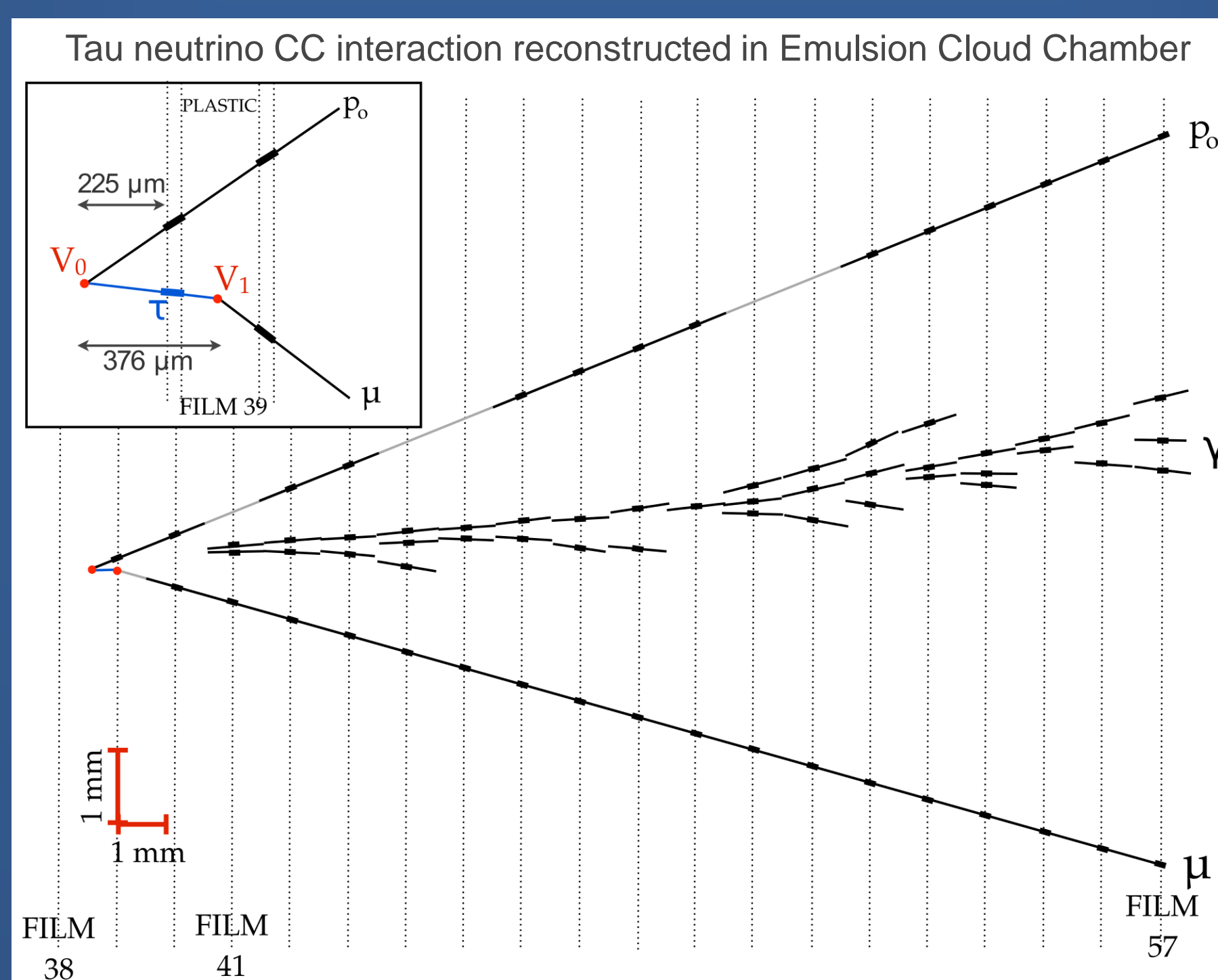


Scintillator planes were used to predict the ECC element in which neutrino interaction occurred. Selected ECCs were then extracted, and contained emulsions were scanned by automatic scanning microscopes.



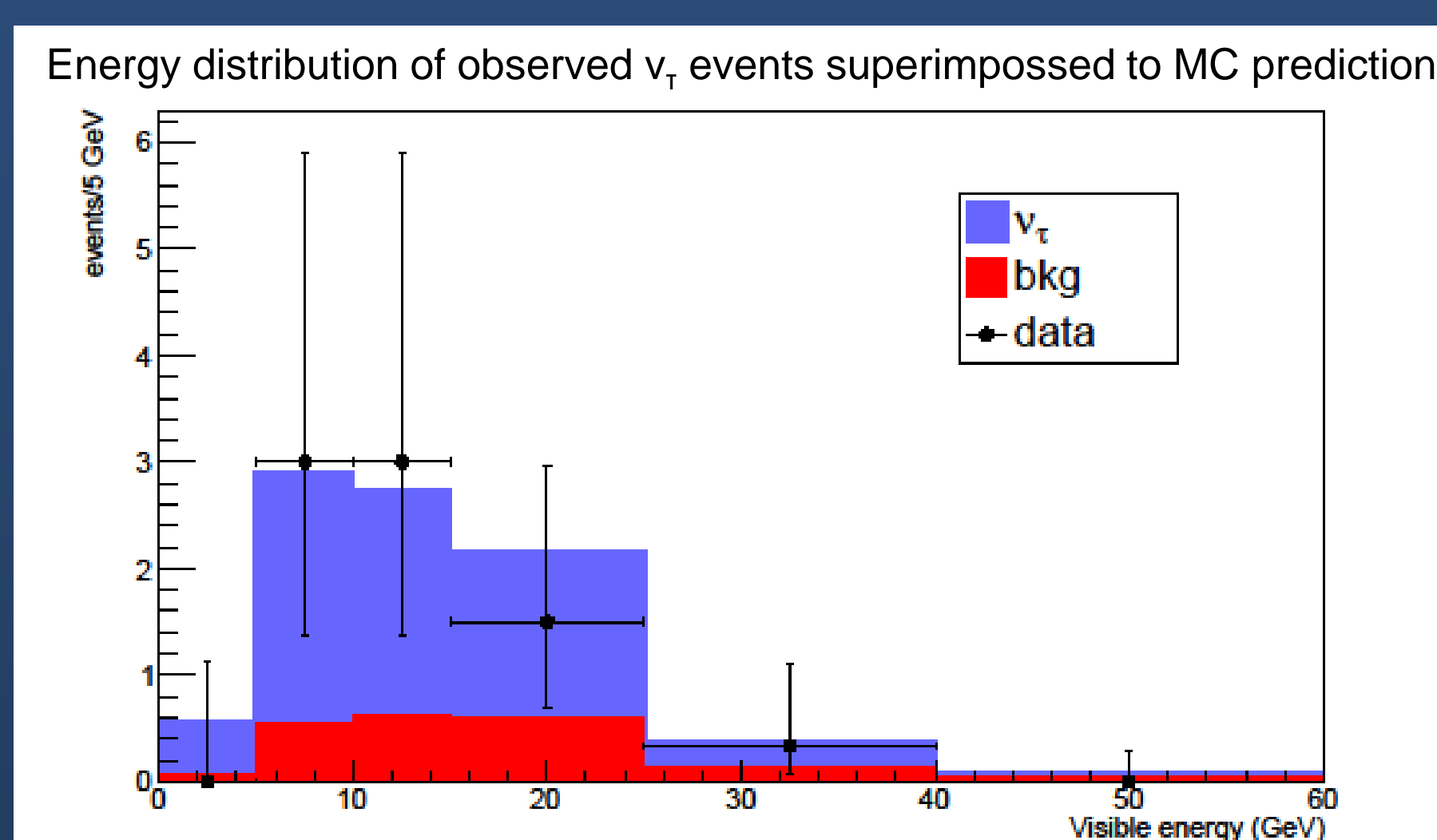
Data taking run lasted from 2008 to 2012 with an exposure of  $17.97 \times 10^{19}$  p.o.t, in which 19505 events were recorded in the target, of which 5603 were fully reconstructed in emulsion.

## The OPERA experiment (cont.)



## $\nu_{\mu} \rightarrow \nu_{\tau}$ appearance channel (std. osc)

Final results on the  $\nu_{\tau}$  appearance channel were published in 2018 [3]. Ten  $\nu_{\tau}$  candidates were observed with expected background of  $2 \pm 0.4$  events, corresponding to significance  $6.1 \sigma$ .

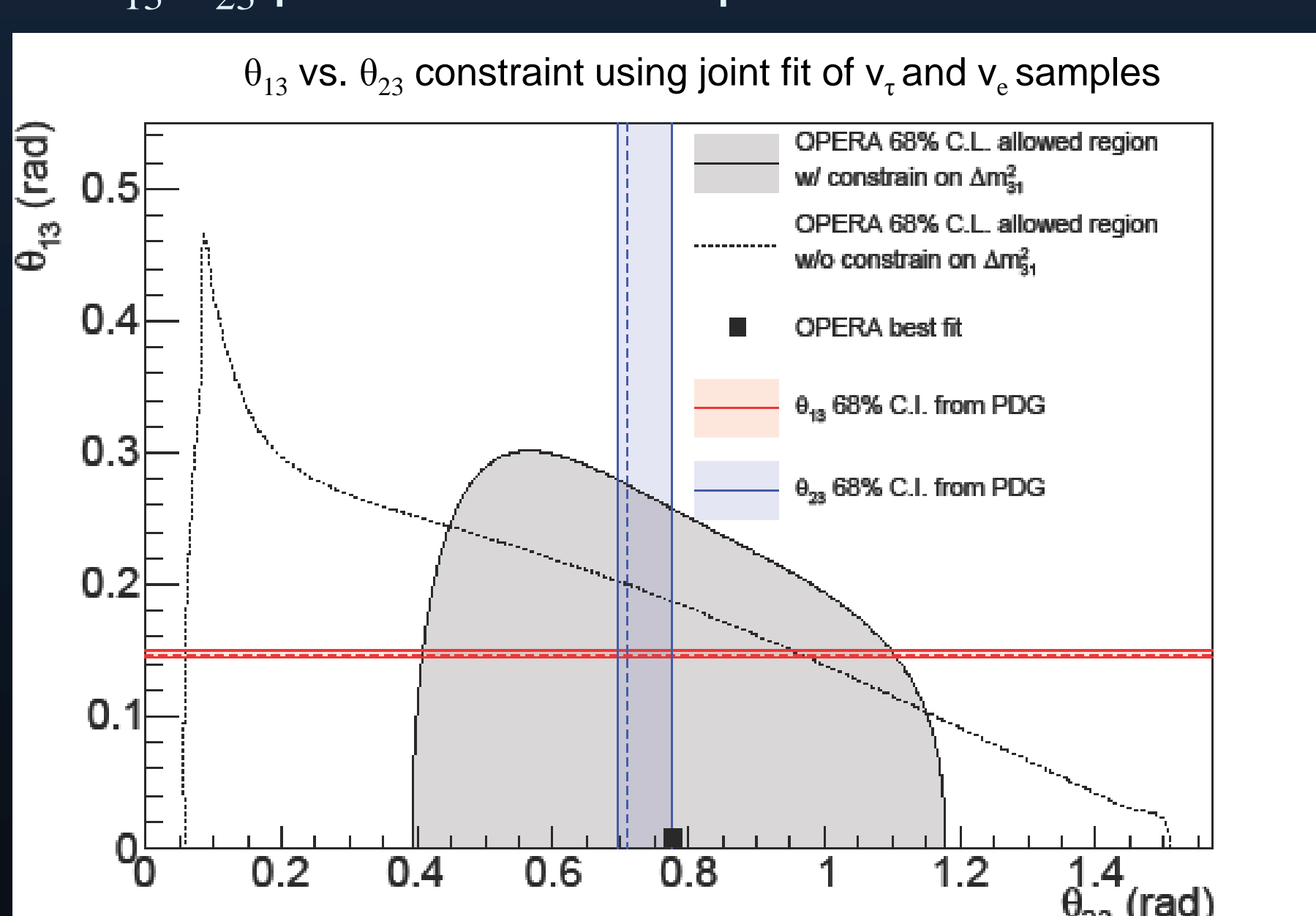


Assuming maximal mixing, OPERA measured the atmospheric mass splitting in  $\nu_{\tau}$  appearance mode:

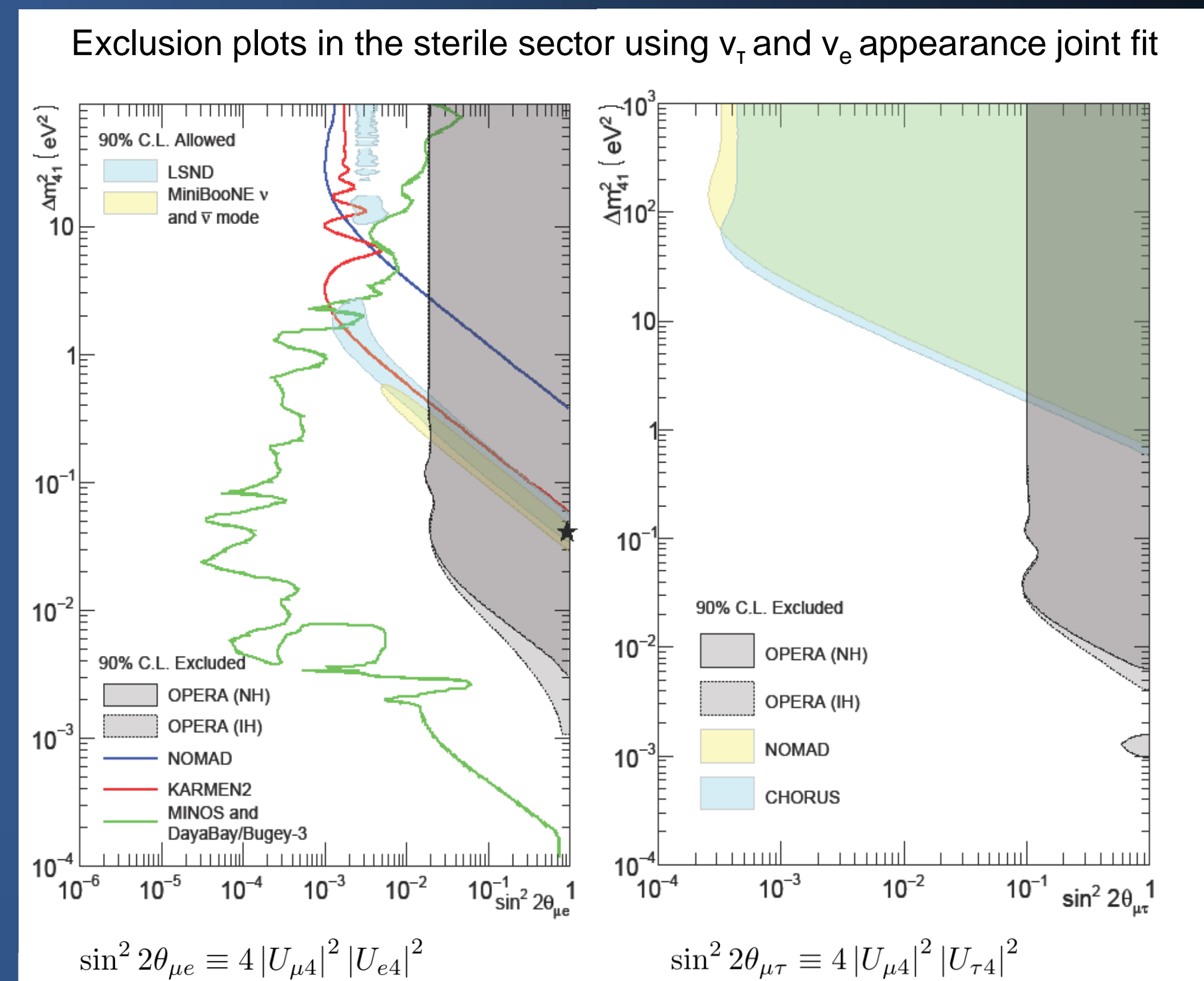
$$|\Delta m_{32}^2| = (2.7_{-0.6}^{+0.7}) \times 10^{-3} \text{ eV}^2$$

## Combined $\nu_{\tau}$ and $\nu_e$ appearance (std. osc.)

OPERA detector is also sensitive to  $\nu_e$  CC interactions. Expected number of  $\nu_e$  candidate events is  $31.9 \pm 3.2$  in case of no oscillations, and  $34.3 \pm 3.4$  for standard oscillations. 35 events were observed, consistent with both cases [4]. A joint fit using  $\nu_{\tau}$  and  $\nu_e$  appearance in  $\theta_{13}$ - $\theta_{23}$  plane has been performed.

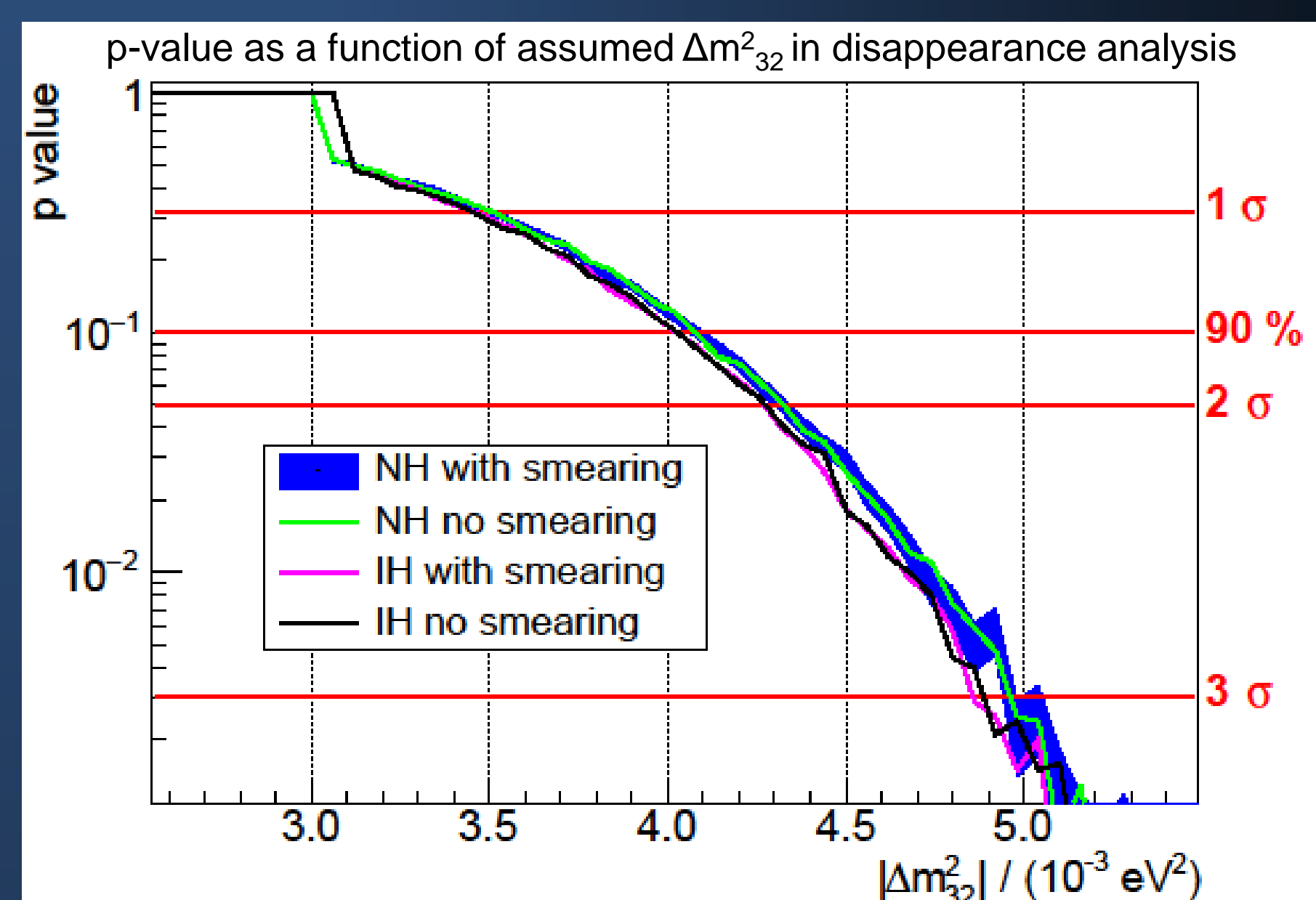


## Combined $\nu_{\tau}$ and $\nu_e$ appearance (sterile)



In the sterile sector, OPERA excludes recent MiniBooNE result [5] with significance of  $3.3 \sigma$ , using joint fit of  $\nu_{\tau}$  and  $\nu_e$  samples.

## $\nu_{\mu}$ disappearance channel



Due to the lack of a near detector, NC/CC event rate ratio was used to search for  $\nu_{\mu}$  disappearance signal. Assuming all other oscillation parameters fixed, an upper limit on atmospheric mass splitting was obtained:

$$|\Delta m_{32}^2| < 4.1 \times 10^{-3} \text{ eV}^2 @ 90\% \text{ C.L.}$$

## Conclusions

OPERA is the only experiment able to study all three channels of  $\nu_{\mu}$  oscillations:  $\nu_{\mu} \rightarrow \nu_{\tau}$ ,  $\nu_{\mu} \rightarrow \nu_e$ ,  $\nu_{\mu} \rightarrow \nu_{\mu}$ . Apart from the primary  $\nu_{\tau}$  analysis, a joint fit of  $\nu_{\tau}$  and  $\nu_e$  has been performed in both standard and sterile hypothesis. Additionally, a search to  $\nu_{\mu}$  disappearance has been performed.

## References

1. R. Acquafredda et al. (OPERA), JINST 4 (2009) P04018
2. K. Elsener, Reports No. CERN 98-02 and No. INFN/AE-98/05, 1998; R. Bailey et al., Reports No. CERN-SL-99-034-DI and No. INFN-AE-99-05, 1999, addendum to Reports No. CERN 98-02 and No. INFN-AE-98-05, 1998; CNGS Web site, <http://proj-cngs.web.cern.ch/proj-cngs>
3. N. Agafonova et al. (OPERA), Phys. Rev. Lett. 120.21 (2018), p. 211801
4. N. Agafonova et al. (OPERA), JHEP 06 (2018), p. 151
5. A. A. Aguilar-Arevalo et al. (MiniBooNE), Phys. Rev. Lett. 121 (2018), p. 221801

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